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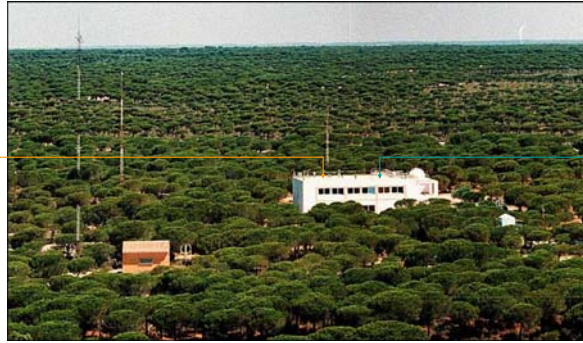
Calibration campaign: from 30th September to 27th October 2005

at the Atmospheric Sounding Station (ESAt/ INTA/ "El Arenosillo" (Huelva, Spain)

INSTRUMENTS AND LOCATION



Three broadband radiometers
UV-S-E-T model (erythema weighted)
Kipp & Zonen



ESAt/ INTA/ "El Arenosillo"
37.1° N, 6.7° W, 20 m a.s.l.



Brewer MK III double monochromator
spectrophotometer (Kipp & Zonen)

CALIBRATION METHODS

ONE STEP METHODS

- ratio model $C_r = \frac{1}{N} \sum_{i=1}^N \frac{UVER_i^{Brewer-CIE}}{V_i}$
- first-order model $UVER_i^{Brewer-CIE} = C_f V_i$
- second-order model $UVER_i^{Brewer-CIE} = C_{s1} V_i + C_{s2} V_i^2$
- **angular method** $UVER_i^{Brewer-CIE} = C_1 V_i + C_2 V_i^2 \cos \theta_i$

$UVER^{Brewer-CIE}$ represents Brewer erythemal irradiance, V the output voltage of broadband radiometer; C_f , C_{s1} , C_{s2} , C_1 and C_2 the regression coefficients and θ the solar zenith angle.

TWO STEPS METHOD

- Calibration matrix $f([O_3], \theta)$ determined using:
- spectral response function of broadband instrument (SRF), applied to spectroradiometer data to obtain an absolute calibration factor
 - transfer radiation model (UVSpec) to obtain spectral UV irradiance (CIE and SRF weighted)

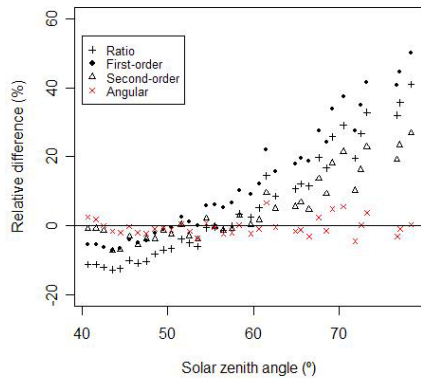
$$UVER^{CIE} = (V - V_o) \cdot C \cdot f_n(\theta, [O_3]) \cdot COSCOR(\theta)$$

Data used
733 minute average data

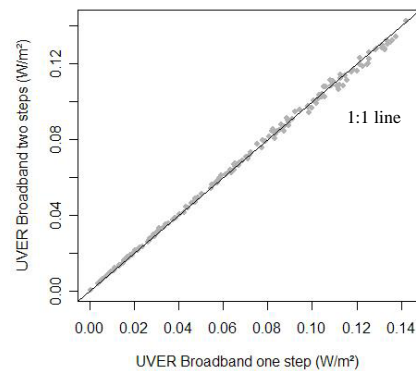
To fit the models:
subset of 564 data randomly selected
(77% of the whole data set)

Validation of the models:
subset of 169 data
(23% remaining)

RESULTS



Relative differences between UV irradiance measured by Brewer spectrophotometer and estimated by broadband radiometer #1 using the four one-step models



Comparison of broadband values obtained applying one-step (angular method) and two-steps methods (matrix calibration method) using subset of 169 data reserved for validation

The new calibration model proposed - **angular method** - means a remarkable improvement of broadband radiometer's calibration procedures based in one-step methodology. The excellent agreement with results obtained by the two-steps method, confirms the reliability and usefulness of the new one-step method proposed, even for high solar zenith angles.

ACKNOWLEDGEMENTS

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