

## Protocol of the intercomparison at ARPA, Aosta, Italy in June-July 2025 with the travelling reference spectroradiometer QASUME from PMOD/WRC

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The purpose of the visit was the comparison of global solar irradiance measurements between the spectroradiometer AAO operated by the Sezione Agenti Fisici - Radiazione Ultravioletta Solare, Agenzia Regionale per la Protezione dell'Ambiente (ARPA) and the travel reference spectroradiometer QASUME. The measurement site is located at Valle d'Aosta; Latitude 45.74 N, Longitude 7.34 E and altitude 569 m.a.s.l.

The horizon of the measurement site is free down to at least 80° solar zenith angle (SZA). Measurements between 5:00 UT and 17:20 UT have been analysed.

QASUME was installed at ARPA Aosta 30 June 2025. The spectroradiometer was installed next to AAO with the entrance optic of QASUME within 1 m of AAO. The spectroradiometer in use at ARPA Aosta is a Bentham DTMc300 double monochromator. The intercomparison between QASUME and the ARPA spectroradiometer lasted 5 days, from the afternoon of June 30 to the morning of July 4, 2025.

QASUME was calibrated several times during the intercomparison period using a portable calibration system. Four lamps (T68523, T61251, T153061 and T153062) were used to obtain a spectral irradiance calibration traceable to the primary reference held at PMOD/WRC, which is traceable to PTB. The daily mean responsivity of the instrument based on these calibrations varied by less than 1 %. The internal temperature of QASUME was  $31.5 \pm 0.4$  °C and the diffuser head was heated to a temperature of  $29.2 \pm 1.2$  °C.

The wavelength shifts relative to the QASUMEFTS (Gröbner et al., 2017) spectrum as retrieved from the matSHIC analysis were between  $\pm 50$  pm in the spectral range 290 to 500 nm.

**Protocol:**

The measurement protocol was to measure one solar irradiance spectrum every 30 minutes from 290 to 500 nm, every 0.25 nm, and 1.5 second between each wavelength increment.

| DOY | Date   | DAY       | Weather                                                                            | Comment (times are in UT)                                                                                        |
|-----|--------|-----------|------------------------------------------------------------------------------------|------------------------------------------------------------------------------------------------------------------|
| 181 | 30-Jun | Monday    | Mostly diffuse sky                                                                 | Installed at 7:15<br>13:27 calibration T153062                                                                   |
| 182 | 01-Jul | Tuesday   | Clear sky in the morning<br>14:00 Cumulus Cloud passing<br>Diffuse sky after 16:00 | 13:26 Calibration T153061<br>14:24 Calibration T61251                                                            |
| 183 | 02-Jul | Wednesday | Diffuse sky<br>11-14:30 thinner cloud layer<br>with partially direct sun           | 09:45 Calibration T153061<br>09:45 Calibration Check AAO<br>10:05 Calibration T68523<br>10:36 Calibration T61251 |
| 184 | 03-Jul | Thursday  | Mix of sun and clouds<br>No direct sun from 12-16:40                               | 09:25 Calibration T68523<br>09:45 Calibration T153062<br>15:59 AAO Head: Rotation by +180°                       |
| 185 | 04-Jul | Friday    | Clear sky in the morning<br>Cumulus clouds after 8:30                              | 08:25 AAO Head: Rotation by -180°<br>09:25 Calibration T153062<br>10:00 QASUME OFF                               |

**Results:**

In total 114 synchronised simultaneous solar spectra from QASUME and AAO are available from the measurement period. Measurements between 5:00 UT and 17:20 UT have been analysed (SZA smaller than 80°).

The spectra from AAO were corrected for wavelength shift and convolved with a 1 nm triangular slit function before being submitted for the calibration.

**Conclusions:**

1. The spectral ratios between AAO and QASUME are on average at unity with a diurnal variability of less than  $\pm 3\%$ .
2. During the scan with direct solar irradiance in the morning a small cosine error effect is revealed through the comparison to QASUME. This can also be seen in the change from direct to total diffuse irradiance at 14:00 UT on day 182. A clear-sky cosine correction to the AAO data was attempted, but it didn't improve the comparison.
3. The rotation test of the AAO input optic from 16:00 (183) to 8:25 (184) didn't reveal any azimuthal asymmetry of the input optic.

**Summary:**

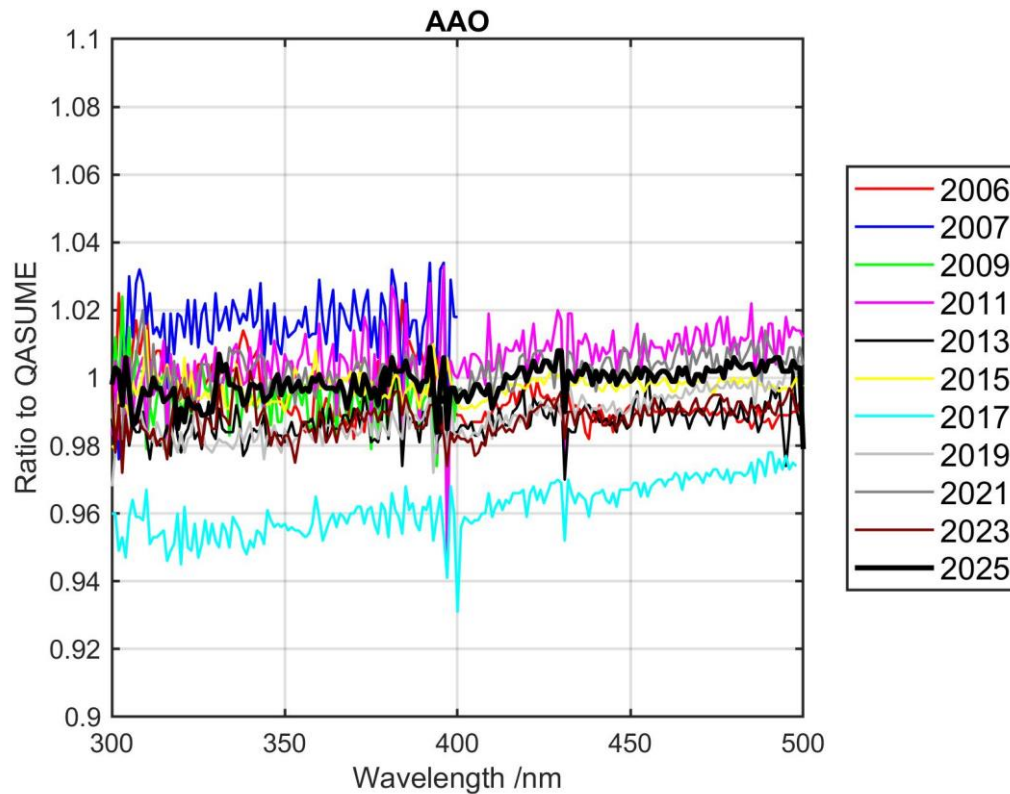
The audit showed that the UV measurements performed at ARPA-Aosta with the AAO spectroradiometer are within the combined expanded uncertainty:

$$U_{\text{comp}} = \sqrt{U_{\text{qas}}^2 + U_{\text{AAO}}^2}$$

With  $U_{\text{qas}} = 1.74\%$  and  $U_{\text{qas}} = 7.34\%$  ( $\lambda < 310 \text{ nm}$ ) (Hülse et al., 2016).  $U_{\text{AAO}}$  was estimated to be 4%.

### **Comparison to previous QASUME site visits**

The stability of AAO was assessed by comparing QASUME visits performed since 2006. As seen in figure 1, the campaign average ratio to QASUME has been stable to  $\pm 2\%$ , except for the comparison performed in 2017 (the issue was explained in detail by Fountoulakis et al., 2020, ESSD and AAO the spectra were post-corrected accordingly).



**Figure 1** solar spectral ratios of AAO to QASUME averaged over each QASUME site visit.

### **References:**

Fountoulakis, I., Diémoz, H., Siani, A. M., Hülsen, G., and **Gröbner, J.**: Monitoring of solar spectral ultraviolet irradiance in Aosta, Italy, *Earth Syst. Sci. Data*, 12, 2787–2810, <https://doi.org/10.5194/essd-12-2787-2020>, 2020.

J. Gröbner, I. Kröger, L. Egli, G. Hülsen, S. Riechelmann, and P. Sperfeld, "The high-resolution extraterrestrial solar spectrum (QASUMEFTS) determined from ground-based solar irradiance measurements", *Atmos. Meas. Tech.*, 10, 3375-3383, 2017

G. Hülsen, J. Gröbner, S. Nevas, P. Sperfeld, L. Egli, G. Porrovecchio, and M. Smid, "Traceability of solar UV measurements using the Qasume reference spectroradiometer", *Appl. Opt.* 55, 7265-7275, 2016.

### **Comments from the operator:**

The irradiance of the final submitted spectra was multiplied by a correction factor of  $1/0.991$  (where 0.991 is the average diffuse angular error over the four planes).

The stability of the AAO instrument during the audit was assessed on July 2 using two 200W lamps recently calibrated against 1000W lamps traceable to PMOD standards. The difference between responses obtained with the two lamps was 0.26%. Comparison with the previous calibration (June 27, 2025) showed -0.36%, confirming instrument stability. The AAO response was updated at 11:30 UT.

## **Appendix**

Detailed results for all local spectrophotometers with respect to the reference spectroradiometer QASUME

